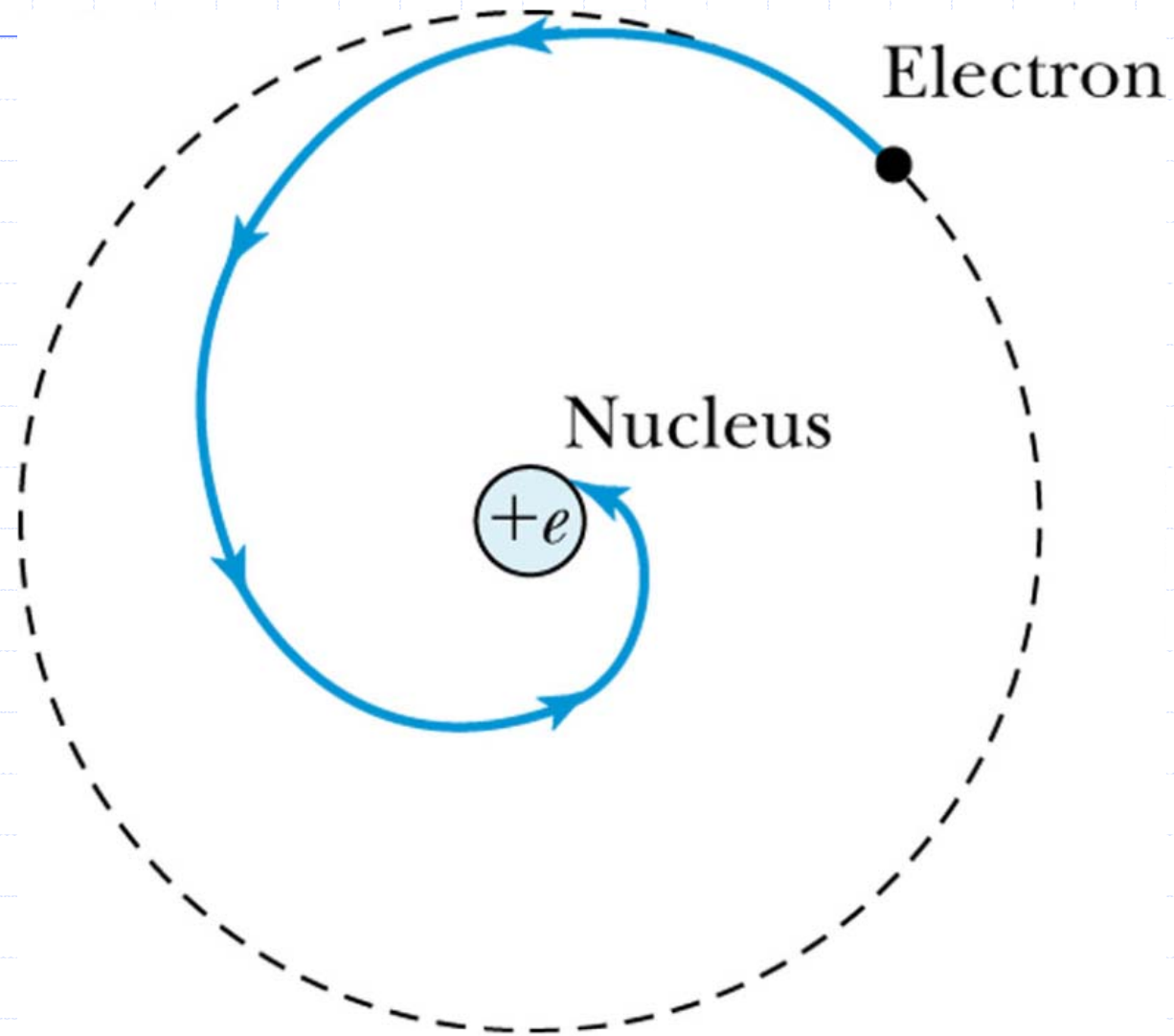


Bohr Model

- Rutherford established the atomic nucleus as a positive charge of radius $\sim 1\text{F}$
- At the same time, the radius of an atom was known to be $\sim 10^{-10}\text{m}$
- A natural model for the atom is a planetary model
- However electrons orbiting the nucleus are accelerating and hence radiate energy as electromagnetic waves

Bohr Model



Bohr Model

➤ Postulates of the Bohr model

- An electron moves in a circular orbit about the nucleus acted on only by the Coulomb force and following classical mechanics

- The only orbits possible are those whose angular momentum

$$L = \frac{nh}{2\pi} = n\hbar$$

- The orbiting electron does not emit electromagnetic radiation
- Electromagnetic radiation is emitted if an electron moves from one orbit to another. The frequency of the radiation is

$$f = \frac{(E_i - E_f)}{h}$$

Bohr Model

- Thus the Bohr model is a mixture of both classical and non-classical physics
- It described known atomic spectra precisely
- It had predictive power
- The math was easy

Bohr Model

- Assume initially that the nucleus is infinitely heavy
- Since the orbit of the electron is stable

$$k \frac{Ze^2}{r^2} = \frac{mV^2}{r}$$

- The orbital angular momentum L is a constant

$$L = mVr = n\hbar, n = 1, 2, 3, \dots$$

$$V = \frac{n\hbar}{mr}$$

Bohr Model

➤ Then

$$kZe^2 = mrV^2 = \frac{mrvn^2\hbar^2}{m^2r^2} = \frac{n^2\hbar^2}{mr}$$

➤ Thus

$$r = \frac{n^2\hbar^2}{kmZe^2} = \frac{n^2a_0}{Z}$$

where the first Bohr radius $a_0 \equiv \frac{4\pi\epsilon_0\hbar^2}{me^2} = 0.53 \times 10^{-10} m$

$$V = \frac{n\hbar}{mr}$$

Bohr Model

➤ Quantization of angular momentum restricts the possible orbits of the electron

- For $Z=1$ and $n=1$ we find $r=0.53 \times 10^{-10} \text{m}$, in good agreement with the known atomic radius

➤ Evaluating the orbital velocity

- For $Z=1$ and $n=1$, we find $v=2.2 \times 10^6 \text{m/s}$
- Relativistic mechanics is not needed (for small Z)

Bohr Model

➤ The total energy of the electron $E = T + V$

$$V = k \int_{\infty}^r \frac{Ze^2}{r^2} dr = -k \frac{Ze^2}{r}$$

$$T = \frac{1}{2} mv^2 = k \frac{Ze^2}{2r}$$

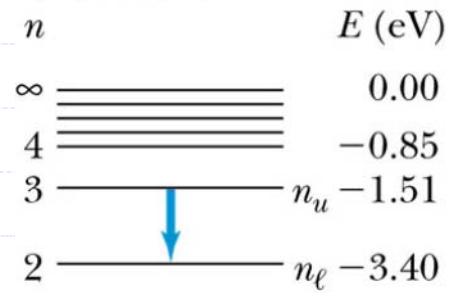
$$E = T + V = -k \frac{Ze^2}{2r} = -T$$

$$E_n = -\frac{Z^2 e^2}{8\pi\epsilon_0 a_0 n^2} \equiv -\frac{Z^2 E_0}{n^2}$$

$$\text{where } E_0 = \frac{e^2}{8\pi\epsilon_0 a_0} = \frac{me^4}{2\hbar^2 (4\pi\epsilon_0)^2} = 13.6 \text{ eV}$$

Bohr Model

➤ Energy levels in hydrogen



↑
Energy

1 ————— -13.6

Bohr Model

- According to Bohr's postulate, emission of electromagnetic radiation occurs when an electron in a higher energy state E_i decays to a lower energy state E_f
- The frequency and wavelength of the emitted radiation is

$$hf = E_i - E_f$$

$$\frac{1}{\lambda} = \frac{E_i - E_f}{hc} = \frac{Z^2 E_0}{hc} \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \equiv Z^2 R_\infty \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

Line Spectra

Hydrogen Series of Spectral Lines

Discoverer (year)	Wavelength	n	k
Lyman (1916)	Ultraviolet	1	>1
Balmer (1885)	Visible, ultraviolet	2	>2
Paschen (1908)	Infrared	3	>3
Brackett (1922)	Infrared	4	>4
Pfund (1924)	Infrared	5	>5

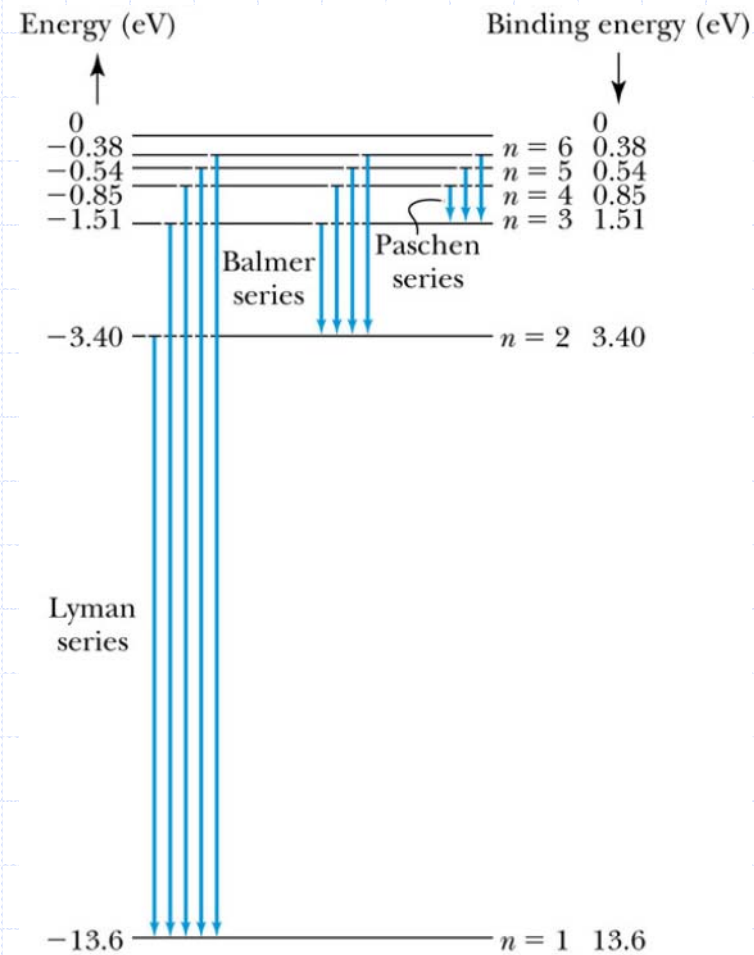
$$\frac{1}{\lambda} = R_H \left(\frac{1}{n^2} - \frac{1}{k^2} \right), \quad R_H = 1.096776 m^{-1}$$

Bohr Model

- Thus the Bohr model correctly gives the wavelengths of (hydrogen) emission and absorption spectra
- Bohr's theory predicted the Lyman, Brackett and Pfund series before they were discovered
- It also predicted the spectra for He^+ before it was observed
- R_∞ depends only on fundamental constants (e, m_e, c, \hbar)
- R_∞ agrees well with R_H (1.097373 vs $1.096776 \times 10^7 \text{m}^{-1}$)

Bohr Model

➤ Radiative transitions in hydrogen



Bohr Model

➤ Suppose a He atom in the ground state absorbs a photon with $\lambda=41.3\text{nm}$. Will the He atom be ionized?

$$E = \frac{hc}{\lambda} = \frac{1240\text{eV} \cdot \text{nm}}{41.3\text{nm}} = 30\text{eV}$$

$$E_1(\text{He}) = \frac{Z^2 E_0}{1^2} = (4)(13.6\text{eV}) = 54.4\text{eV}$$

Bohr Model

➤ We assumed the nucleus was infinitely heavy but in fact the electron and nucleus revolve around their mutual center-of-mass

- If r is the distance between electron and nucleus

$$m(r - x) = Mx$$

$$x = \frac{mr}{m + M} \text{ and } r - x = \frac{Mr}{m + M}$$

- The total angular momentum of the system is

$$L = m(r - x)^2 \omega + Mr^2 \omega$$

$$L = \frac{mM^2 r^2 \omega}{(m + M)^2} + \frac{Mm^2 r^2 \omega}{(m + M)^2}$$

$$L = \frac{mM}{m + M} r^2 \omega$$

Bohr Model

➤ Summarizing

$$L = \frac{mM}{m + M} r^2 \omega$$

■ Check. As $M \rightarrow \infty$, $L \rightarrow mr^2\omega$ and $x \rightarrow 0$

➤ Following Bohr's postulate we let

$$L = \frac{mM}{m + M} r^2 \omega = \frac{mM}{m + M} Vr = n\hbar$$

➤ Our previous calculations then follow with the substitution

$$m \rightarrow \frac{mM}{m + M} \equiv \text{reduced mass } \mu$$

Bohr Model

➤ Thus

$$E_n = -\frac{Z^2 e^2}{8\pi\epsilon_0 a_0 n^2} \equiv -\frac{Z^2 E_0}{n^2}$$

$$\text{where } E_0 = \frac{e^2}{8\pi\epsilon_0 a_0} = \frac{\mu e^4}{2\hbar^2 (4\pi\epsilon_0)^2} = 13.6\text{eV}$$

$$\frac{1}{\lambda} = \frac{E_i - E_f}{hc} = \frac{Z^2 E_0}{hc} \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \equiv Z^2 R \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

➤ So $R = \left(\frac{\mu}{m} \right) R_\infty$

➤ And now R and R_H agree to within 3 parts in 100000

Bohr Model

➤ Calculate the three lowest energy states of positronium

$$\mu = \frac{m_e m_e}{m_e + m_e} = \frac{m_e}{2}$$

so

$$E_n = -\frac{Z^2 E_0}{n^2} = -\frac{Z^2 m e^4}{2\hbar^2 (4\pi\epsilon_0)^2}$$

becomes

$$E_n = -\frac{Z^2 E_0}{2n^2}$$

$$E_1 = -6.8eV, E_2 = -1.7eV, E_3 = -0.76eV$$

Bohr Model

➤ While the Bohr model simply and successfully described some of the features of atoms, it failed on other aspects

- It was not applicable to multi-electron atoms
- It was not able to predict the transition rate between atomic states
- It could not account for the fine structure observed in spectral lines
- It could explain how atoms bind into molecules
- It was an ad-hoc mixture of ideas